## CS-I 84: Computer Graphics

Lecture \#|7: Global Illumination

Prof. James O'Brien
University of California, Berkeley
vol| $1 \cdot 17.10$ $\qquad$

Today

- The Rendering Equation
- Radiosity Method
- Photon Mapping
- Ambient Occlusion



## The Rendering Equation


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## The Rendering Equation



The light shining on $\times$ from $x^{\prime}$ is equal to:

- the emitted light from $x$ toward $x$, plus
- for each bit of surface in the scene, how much light shines from that bit onto $x$ ' and is reflected toward $x$, scaled appropriately
$L_{s}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\delta\left(\mathbf{x}, \mathbf{x}^{\prime}\right)\left[E\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+\int_{S} \rho_{x^{\prime}}\left(\mathbf{x}, \mathbf{x}^{\prime \prime}\right) L_{s}\left(\mathbf{x}^{\prime}, \mathbf{x}^{\prime \prime}\right) \frac{\cos \left(\theta^{\prime}\right) \cos \left(\theta^{\prime \prime}\right)}{\left\|\mathbf{x}^{\prime}-\mathbf{x}^{\prime \prime}\right\| \|^{2}} \mathrm{~d} \mathbf{x}^{\prime \prime}\right]$

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## The Rendering Equation

$L_{s}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\delta\left(\mathbf{x}, \mathbf{x}^{\prime}\right)\left[E\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+\int_{S} \rho_{x^{\prime}}\left(\mathbf{x}, \mathbf{x}^{\prime \prime}\right) L_{s}\left(\mathbf{x}^{\prime}, \mathbf{x}^{\prime \prime}\right) \frac{\cos \left(\theta^{\prime}\right) \cos \left(\theta^{\prime \prime}\right)}{\left\|\mathbf{x}^{\prime}-\mathbf{x}^{\prime \prime}\right\|^{2}} \mathrm{~d} \mathbf{x}^{\prime \prime}\right]$ $\qquad$

| The Rendering Equation |
| :---: |
| $L_{s}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\delta\left(\mathbf{x}, \mathbf{x}^{\prime}\right)\left[E\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+\int_{S} \rho_{s^{\prime}\left(\mathbf{x}, \mathbf{x}^{\prime \prime}\right) L_{s}\left(\mathbf{x}^{\prime}, \mathbf{x}^{\prime \prime}\right) \frac{\cos \left(\theta^{\prime}\right) \cos \left(\theta^{\prime \prime}\right)}{\left\\|\mathbf{x}^{\prime}-\mathbf{x}^{\prime \prime}\right\\|^{2}} \mathrm{~d} \mathbf{x}^{\prime \prime}}^{\text {Light energy hitting } \times \text { from } x^{\prime}}\right.$ |
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The Rendering Equation

$$
L_{s}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\delta\left(\mathbf{x}, \mathbf{x}^{\prime}\right)\left[E\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+\int_{S} \rho_{x}\left(\mathbf{x}, \mathbf{x}^{\prime \prime}\right) L_{s}\left(\mathbf{x}^{\prime}, \mathbf{x}^{\prime \prime}\right) \frac{\cos \left(\theta^{\prime}\right) \cos \left(\theta^{\prime \prime}\right)}{\left\|\mathbf{x}^{\prime}-\mathbf{x}^{\prime \prime}\right\|^{2}} \mathrm{~d} \mathbf{x}^{\prime \prime}\right]
$$



The Rendering Equation
$L_{s}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\delta\left(\mathbf{x}, \mathbf{x}^{\prime}\right)\left[E\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+\int_{S} \rho_{x}\left(\mathbf{x}, \mathbf{x}^{\prime \prime}\right) L_{s}\left(\mathbf{x}^{\prime}, \mathbf{x}^{\prime \prime}\right) \frac{\cos \left(\theta^{\prime}\right) \cos \left(\theta^{\prime \prime}\right)}{\left\|\mathbf{x}^{\prime}-\mathbf{x}^{\prime \prime}\right\|^{2}} \mathrm{~d} \mathbf{x}^{\prime \prime}\right]$


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The Rendering Equation

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L_{s}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\delta\left(\mathbf{x}, \mathbf{x}^{\prime}\right)\left[E\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+\int_{S} \rho_{x}\left(\mathbf{x}, \mathbf{x}^{\prime \prime}\right) L_{s}\left(\mathbf{x}^{\prime}, \mathbf{x}^{\prime \prime}\right) \frac{\cos \left(\theta^{\prime}\right) \cos \left(\theta^{\prime \prime}\right)}{\left\|\mathbf{x}^{\prime}-\mathbf{x}^{\prime \prime}\right\|^{2}} \mathrm{~d} \mathbf{x}^{\prime \prime}\right]
$$




## The Rendering Equation <br> $L_{s}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\delta\left(\mathbf{x}, \mathbf{x}^{\prime}\right)\left[E\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+\int_{S} \rho_{x}\left(\mathbf{x}, \mathbf{x}^{\prime \prime}\right) L_{s}\left(\mathbf{x}^{\prime}, \mathbf{x}^{\prime \prime}\right) \frac{\cos \left(\theta^{\prime}\right) \cos \left(\theta^{\prime \prime}\right)}{\left\|\mathbf{x}^{\prime}-\mathbf{x}^{\prime \prime}\right\|^{2}} \mathrm{~d} \mathbf{x}^{\prime \prime}\right]$

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|  | Radiosity |
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| - Assume all materials are perfectly Lambertian (diffuse only, |  |
| no specularities) |  |
| - Removes all dependance on directions |  |
| - Reduces dimensionality of lighfield |  |
| - Allows a FM solution (break up into chunks) |  |
| - Can also relax assumption slightly |  |
|  | $\square$ |




| Assume Lambertian |  |
| :---: | :---: |
| $\begin{aligned} & L_{s}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\delta\left(\mathbf{x}, \mathbf{x}^{\prime}\right)\left[E\left(\mathbf{x}, \mathbf{x}^{\prime}\right)+\int_{S} \rho_{x^{\prime}}\left(\mathbf{x}, \mathbf{x}^{\prime \prime}\right) L_{s}\left(\mathbf{x}^{\prime}, \mathbf{x}^{\prime \prime}\right) \frac{\cos \left(\theta^{\prime}\right) \cos \left(\theta^{\prime \prime}\right)}{\left\\|\mathbf{x}^{\prime}-\mathbf{x}^{\prime \prime}\right\\|^{2}} \mathrm{~d} \mathbf{x}^{\prime \prime}\right. \\ & L_{s}\left(\mathbf{x}, \mathbf{x}^{\prime}\right)=\frac{\delta\left(\mathbf{x}, \mathbf{x}^{\prime}\right)}{}\left[E_{x^{\prime}}+\int_{S} \rho_{x^{\prime}} L_{s}\left(\mathbf{x}^{\prime}, \mathbf{x}^{\prime \prime}\right) \frac{\cos \left(\theta^{\prime}\right) \cos \left(\theta^{\prime \prime}\right)}{\left\\|\mathbf{x}^{\prime}-\mathbf{x}^{\prime \prime}\right\\|^{2}} \mathrm{~d} \mathbf{x}^{\prime \prime}\right] \\ & \text { Only term dependent on } \mathbf{x} \end{aligned}$ |  |
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## Discretize into Patches




Discretize into Patches


The candessicicr Theater
Mank Nacke Acortiects.

## Rewrite in Terms of Patches

$$
\begin{gathered}
H_{x^{\prime}}=E_{x^{\prime}}+\rho_{x^{\prime}} \int_{S} \delta\left(\mathbf{x}^{\prime}, \mathbf{x}^{\prime \prime}\right) \frac{H_{x^{\prime \prime}} \cos \left(\theta^{\prime}\right) \cos \left(\theta^{\prime \prime}\right)}{2 \pi} \frac{\left\|\mathbf{x}^{\prime}-\mathbf{x}^{\prime \prime}\right\|^{2}}{} \mathrm{~d} \mathbf{x}^{\prime \prime} \\
H_{i}=E_{i}+\rho_{i} \sum_{j} H_{j} \int_{S_{j}} \delta_{i j} \frac{\cos \left(\theta_{i}\right) \cos \left(\theta_{j}\right)}{2 \pi\left\|\mid \mathbf{c}_{i}-\mathbf{x}\right\|^{2}} \mathrm{~d} \mathbf{x}
\end{gathered}
$$



## Rewrite in Terms of Patches

$$
\begin{aligned}
& H_{x^{\prime}}=E_{x^{\prime}}+\rho_{x^{\prime}} \int_{S} \delta\left(\mathbf{x}^{\prime}, \mathbf{x}^{\prime \prime}\right) \frac{H_{x^{\prime}} \cos \left(\theta^{\prime}\right) \cos \left(\theta^{\prime \prime}\right)}{2 \pi} \frac{\left\|\mathbf{x}^{\prime}-\mathbf{x}^{\prime \prime}\right\|^{2}}{} \mathrm{~d} \mathbf{x}^{\prime \prime} \\
& H_{i}=E_{i}+\rho_{i} \sum_{j} H
\end{aligned}
$$

$$
\text { Form factor from } \mathrm{j} \text { to } \mathrm{i}, F_{i j} \backslash
$$

## Rewrite in Terms of Patches

$$
\begin{aligned}
& H_{x^{\prime}}=E_{x^{\prime}}+\rho_{x^{\prime}} \int_{S} \delta\left(\mathbf{x}^{\prime}, \mathbf{x}^{\prime \prime}\right) \\
& H_{i}=E_{i}+\rho_{i} \sum_{j} H^{2 \pi} \frac{\cos \left(\theta^{\prime}\right) \cos \left(\theta^{\prime \prime}\right)}{\left\|\mathbf{x}^{\prime}-\mathbf{x}^{\prime \prime}\right\|^{2}} \mathrm{~d} \mathbf{x}^{\prime \prime}
\end{aligned}
$$



Example of a rough approximation:

$$
F_{i j} \approx \delta_{i j} \frac{\cos \left(\theta_{i}\right) \cos \left(\theta_{j}\right)}{2 \pi\left\|\mathbf{c}_{i}-\mathbf{c}_{j}\right\|^{2}} A_{j}
$$

## Radiosity Method

- Given the $E_{i}$ and $\rho_{i}$
- First compute $F_{i j}$

- Comments:
- The matrix $\mathbf{A}$ is typically very large
- It is also sparse (why?)
- Should be solved with an iterative method
- e.g.:Jacobi or Gauss-Seidel

Solution is view independent

## Radiosity Method

- Given the light emitted and surface properties
- First compute $F_{i j}$, form factors between patches
- Then solve a linear system to balance energy between all patches
- Comments:
- The system is very large
- It is also sparse (why?)
- Should be solved with an iterative method
- e.g.: Jacobi or Gauss-Seidel
- Solution is view independent

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Progressive Radiosity

- If magnitude of eigenvalues of $\mathbf{A}<1$

$$
(\mathbf{I}-\mathbf{A})^{-1}=\mathbf{I}+\mathbf{A}+\mathbf{A}^{2}+\mathbf{A}^{3}+\cdots
$$

- True for form-factor matrices
- Use Gauss-Seidel-like iteration but reorder by priority

$$
\begin{aligned}
& \mathbf{h}^{k+1}=\mathbf{h}^{k}+\mathbf{u}^{k+1} \\
& \mathbf{u}^{k+1}=\mathbf{A} \mathbf{u}^{k} \\
& \mathbf{h}^{0}=0 \quad \mathbf{u}^{0}=\mathbf{e}
\end{aligned}
$$

| Progressive Rad | diosity |
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|  | TOUChup |
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|  | Each patch will have a constant color <br> • Smooth solution (e.g. average to vertices) |

## OtherThings

- Each patch will have a constant color
- Smooth solution (e.g. average to vertices)

No specular reflection

- Add Phong specular term or raytraced specular reflection
- Grid artifacts
- Be clever with grid...
$\longrightarrow$
$\qquad$
$\square$

| Hierarchical Radiosity |
| :--- | :--- |
| • Light smoothes with distance |
| $\cdot$ compare $1 / h^{2}$ with $1 /\left(h^{2}+d^{2}\right)$ as $h$ gets large |
| $\square$ |
| $\square$ |

## Hierarchical Radiosity

- Light smoothes with distance
- Compare $1 / h^{2}$ with $1 /\left(h^{2}+d^{2}\right)$ as $h$ gets large

Group patches into hierarchy

- Far interactions use lower-res form factors

$\qquad$


## Computing Form Factors

- Form factors have a geometric meaning


Images from
SIGGRAPH 9
SIGGRAPH 93 Education Slide Set
by Stephen Spencer

## Computing Form Factors

- Form factors have a geometric meaning
- "Hemicube" algorithm uses regular scan conversion

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## Computing Form Factors

- Form factors have a geometric meaning
- "Hemicube" algorithm uses regular scan conversion
- Also computed by ray-based sampling
- In practice, computing form factors is the bottleneck


## Photon Mapping

- Lights cast "photons" into environment
- Cast in random directions
- Trace into environment
- Store records at intersections



## Photon Mapping

- Lights cast "photons" into environment
- Cast in random directions
- Trace into environment
- Store records at intersections
- With KD-Trees...









## Ambient Occlusion

- A "hack" to create more realistic ambient illumination cheaply
- Assume light from everywhere is partially blocked by local objects
- At a point on the surface cast rays at random
- Ambient term is proportional to percent of rays that hit nothing
- Weight average by cosine of angle with normal
- Take into account how far before occluded
Ambient Occlusion

|  | Ambient Occlusion |
| :--- | :--- |
| Diffuse Only Ambient Occlusion Combined |  |

Ambient Occlusion

nVidia Gelato Demo Image

