CS-184: Computer Graphics	
Lecture 16: Radiometry	
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V2011-F-16-1.0	

Today		
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Radiometry: measuring light		
 Local Illumination and Raytracing were discussed in an <i>ad hoc</i> fashion 		
Proper discussion requires proper units		
• Not just pretty pictures but correct pictures		
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Matching Reality



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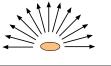
Units	
	-
 Light energy Really power not energy is what we measure 	
 Joules / second (J/s) = Watts (W) 	
 Spectral energy density Power per unit spectrum interval 	
Watts / nano-meter (W/nm)	
Properly done as function over spectrumOften just sampled for RGB	
Often Just sampled for KGB	
Often we assume people know we're talking about S.E.D. and just say E	
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Irradiance
Total light striking surface from all directions
• Only meaningful w.r.t. a surface • Power per square meter (W/m^2) • Really S.E.D. per square meter $(W/m^2/nm)$
• Not all directions sum the same because of foreshortening
* *

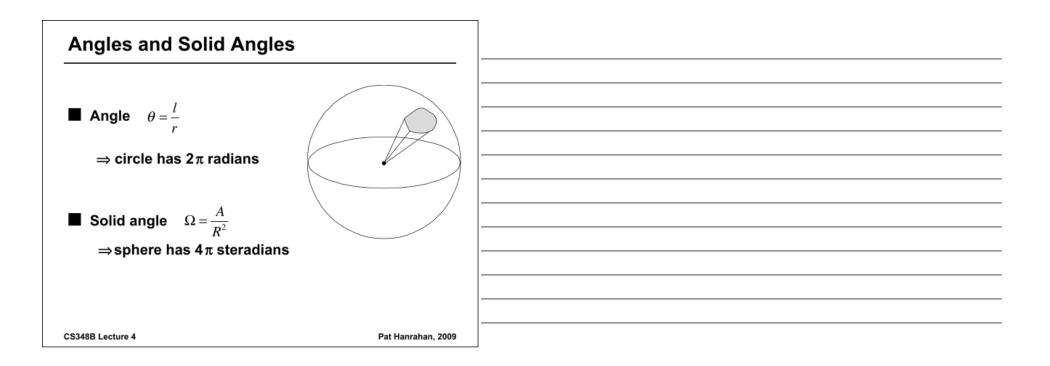
Radiant Exitance

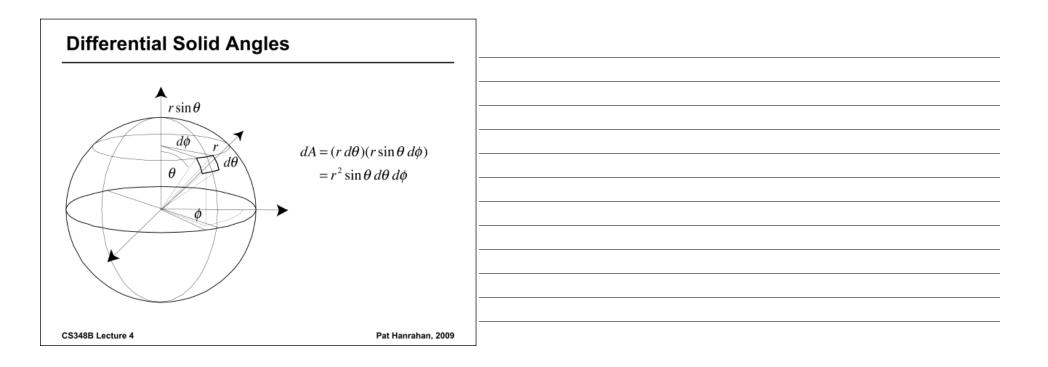
Total light *leaving* surface over all directions

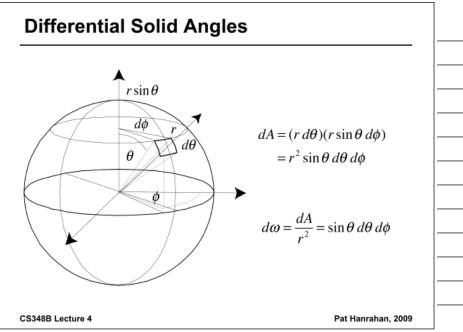
- Only meaningful w.r.t. a surface
- Power per square meter (W/m^2)
- Really S.E.D. per square meter $(W/m^2/nm)$
- Also called Radiosity
- Sum over all directions \Rightarrow same in all directions



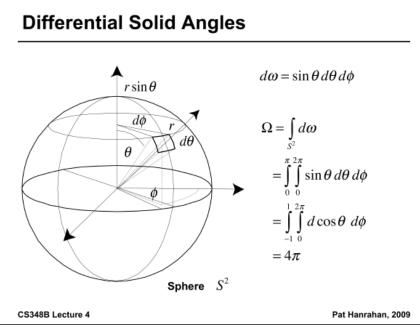
Solid Angles Regular angles measured in radians [0..2π] • Measured by arc-length on unit circle Solid angles measured in steradians [0..4π] • Measured by area on unit sphere • Not necessarily little round pieces...













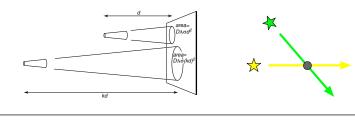
Radiance

Light energy passing though a point in space within a given solid angle

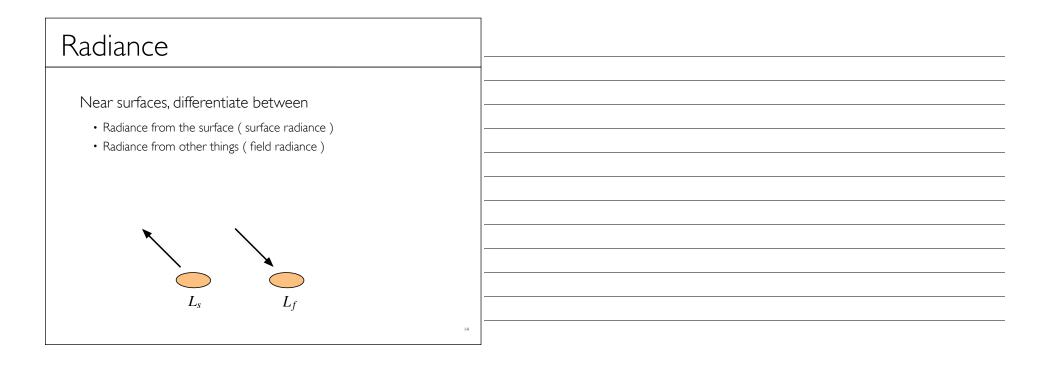
- Energy per steradian per square meter $(W/m^2\,/sr)$
- + S.E.D. per steradian per square meter ($W/m^2/sr/nm$)

Constant along straight lines in free space

 Area of surface being sampled is proportional to distance and light inversely proportional to squared distance

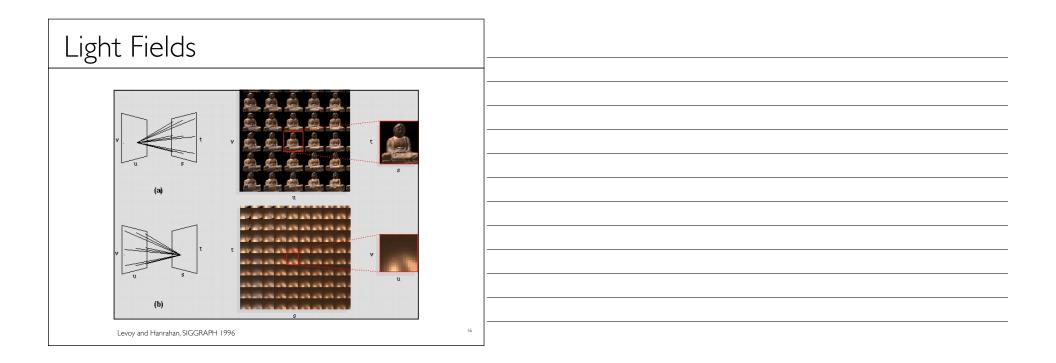






Light Fields

Radiance at every point in space, direction, and frequency: 6D function Collapse frequency to RGB, and assume free space: 4D function Sample and record it over some volume



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Light Fields



Levoy and Hanrahan, SIGGRAPH 1996



Light Fields	
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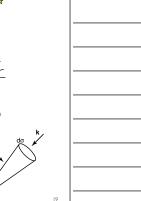
Computing Irradiance

Integrate incoming radiance (field radiance) over all direction

• Take into account foreshortening

$$H = \int_{\Omega} L_f(\mathbf{k}) \cos(\theta) d\sigma$$

$$H = \int_{0}^{2\pi} \int_{0}^{\pi/2} L_f(\theta, \phi) \cos(\theta) \sin(\theta) d\theta d\phi$$



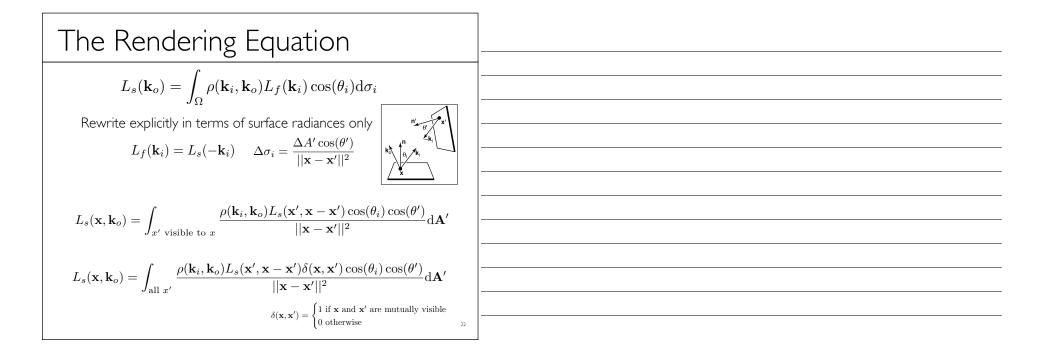
Revisiting The BRDF How much light from direction k_i goes out in direction k_o Now we can talk about units: • BRDF is ratio of surface radiance to the foreshortened field radiance Image: the state of surface radiance to the foreshortened field radiance Image: the state of surface radiance to the foreshortened field radiance Image: the state of surface radiance to the foreshortened field radiance Image: the state of surface radiance to the foreshortened field radiance Image: the state of surface radiance to the foreshortened field radiance Image: the state of surface radiance to the foreshortened field radiance Image: the state of surface radiance to the foreshortened field radiance Image: the state of surface radiance to the state of the sta

The Rendering Equation

Total light going out in some direction is given by an integral over all incoming directions:

$$L_s(\mathbf{k}_o) = \int_{\Omega} \rho(\mathbf{k}_i, \mathbf{k}_o) L_f(\mathbf{k}_i) \cos(\theta_i) \mathrm{d}\sigma_i$$

• Note, this is recursive (my L_f is another's L_s)



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Light Paths

Many paths from light to eye

Characterize by the types of bounces

- Begin at light
- End at eye
- "Specular" bounces
- "Diffuse" bounces

